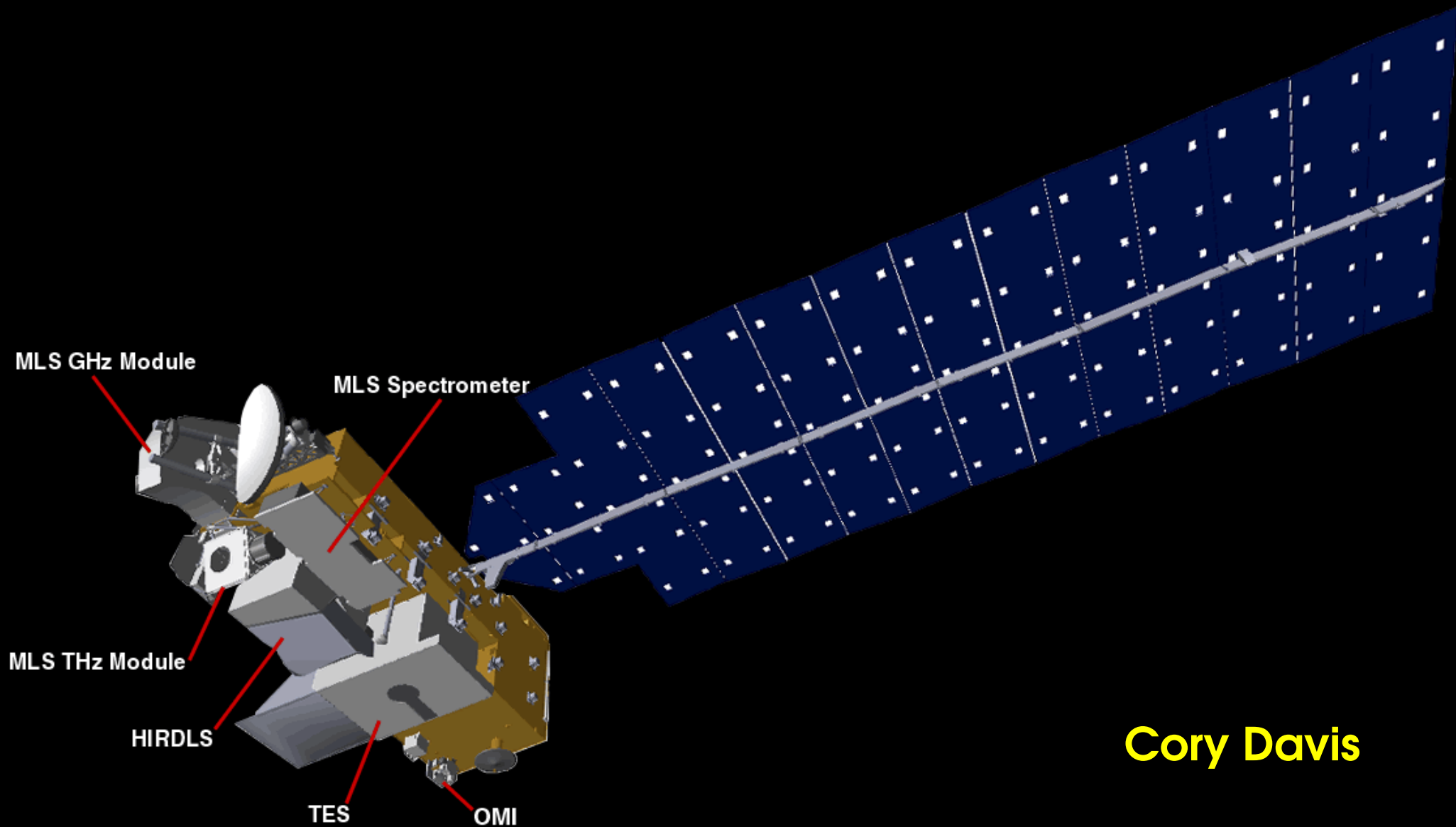
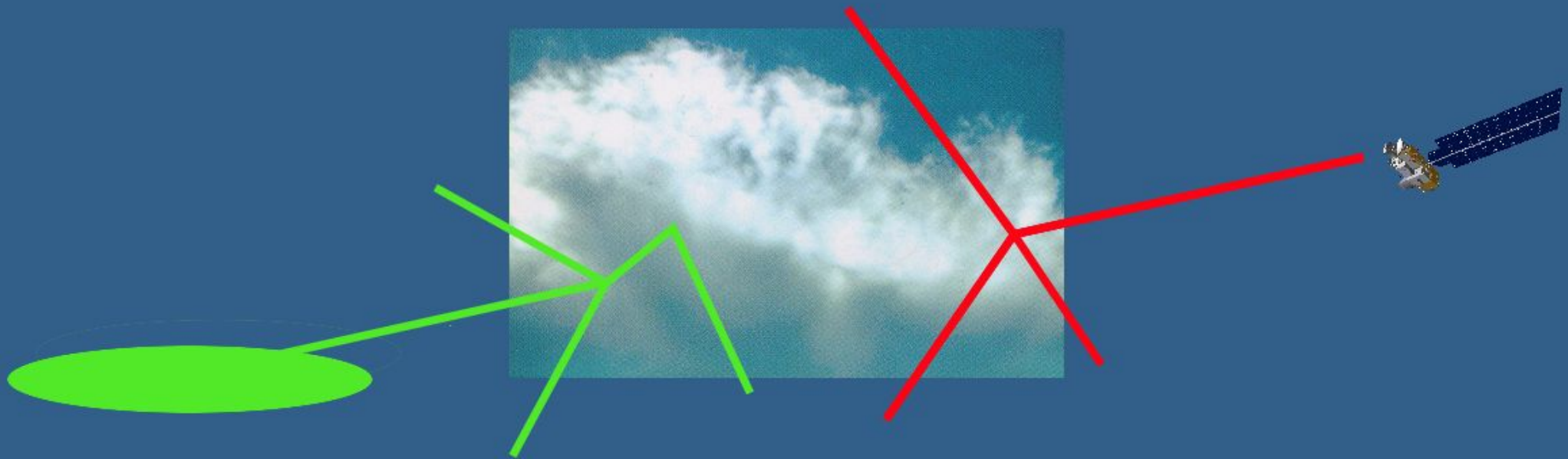


# Aura MLS Cloud Observations



**Cory Davis**

# Clouds and MLS



- Scattering by clouds disrupts composition/ temperature measurements.
- Effect of clouds described by  $\Delta I_{cir} = I_{cloudy} - I_{clear}$
- in optically thin cases this is proportional to IWP

# Polarization

- With scattering comes polarization, defined by the Stokes vector

$$\mathbf{I} = [I, Q, U, V]^T$$

where  $I = I_v + I_h$ ,  $Q = I_v - I_h$

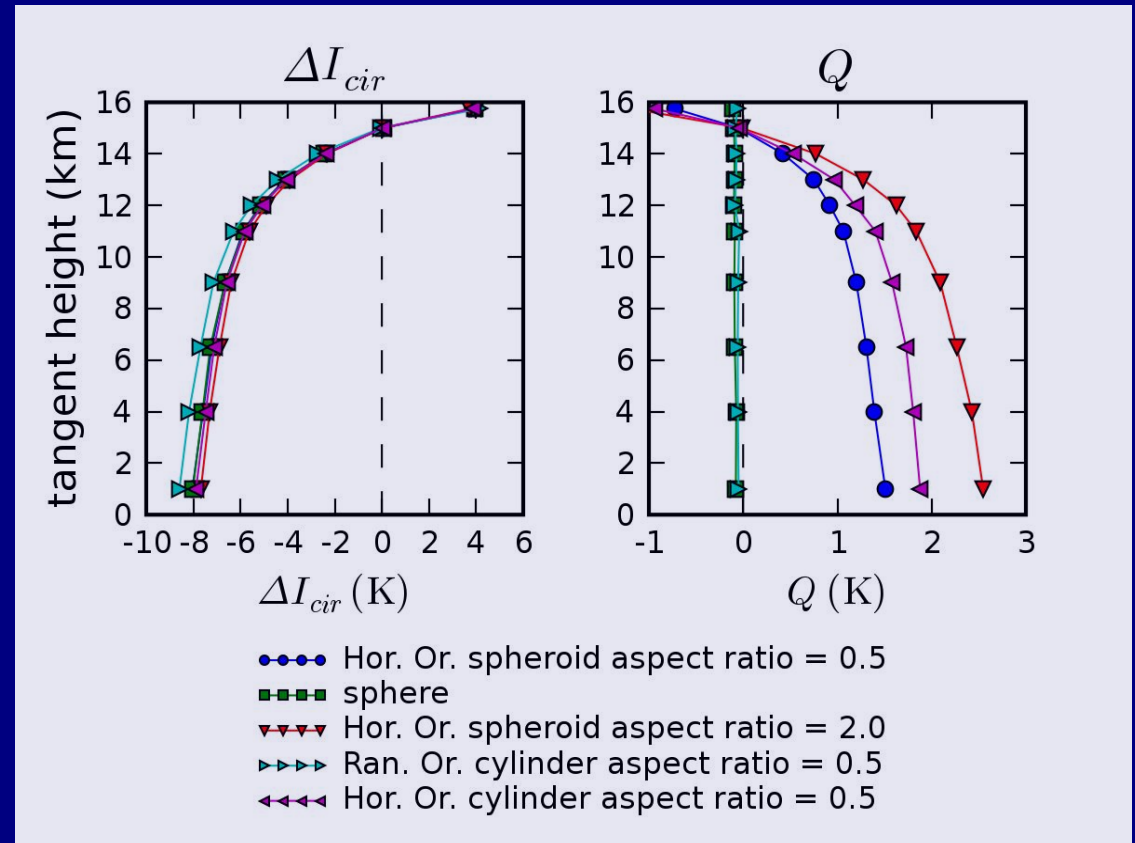
- Simulation studies (with simplified single habit clouds) outline the possibility of significant polarization signal (15 K) for MLS due to preferential horizontal orientation of column and plate crystals.
- Column and plate crystals are generally outnumbered by complex aggregate crystals with unclear orientation behaviour -> the extent to which preferential orientation affects optical properties in real cirrus is unclear

# Aura MLS observations

- First dual polarized microwave limb sounding observations of cirrus
- MLS has five radiometers, two of these measure both H and V polarizations. We focus on Radiometer 1, which is centered on an O<sub>2</sub> line at 118 GHz.
- Not ideal for cloud detection
  - O<sub>2</sub> absorption => only sensitive to high cloud
  - long wavelength => clouds relatively transparent
  - but still can get  $\Delta I_{cir}$  as large as -50 K.

# What simulations tell us to expect

- Horizontally aligned particles give partial vertical polarisation (+ive  $Q$ ) with magnitude decreasing with tangent height
- $I$ ,  $Q$  for horizontal cylinders or prolate spheroids can be reproduced by oblate spheroids with appropriate aspect ratio
- Randomly oriented particles and spheres give negligible polarisation



# Understanding $\mathbf{Q}$

$$\frac{d\mathbf{I}(\mathbf{n})}{ds} = -\mathbf{K}(\mathbf{n})\mathbf{I}(\mathbf{n}) + \mathbf{K}_a(\mathbf{n})I_b(T) + \int_{4\pi} \mathbf{Z}(\mathbf{n}, \mathbf{n}')\mathbf{I}(\mathbf{n}')d\mathbf{n}'$$

small  $\tau$  approximation:

$$Q \approx (J_Q - K_{12}I_0) s$$

$$J_Q = K_{a2}(\mathbf{n})I_b(T) + \int_{4\pi} Z_{2i}(\mathbf{n}, \mathbf{n}')I_i(\mathbf{n}')d\mathbf{n}'$$

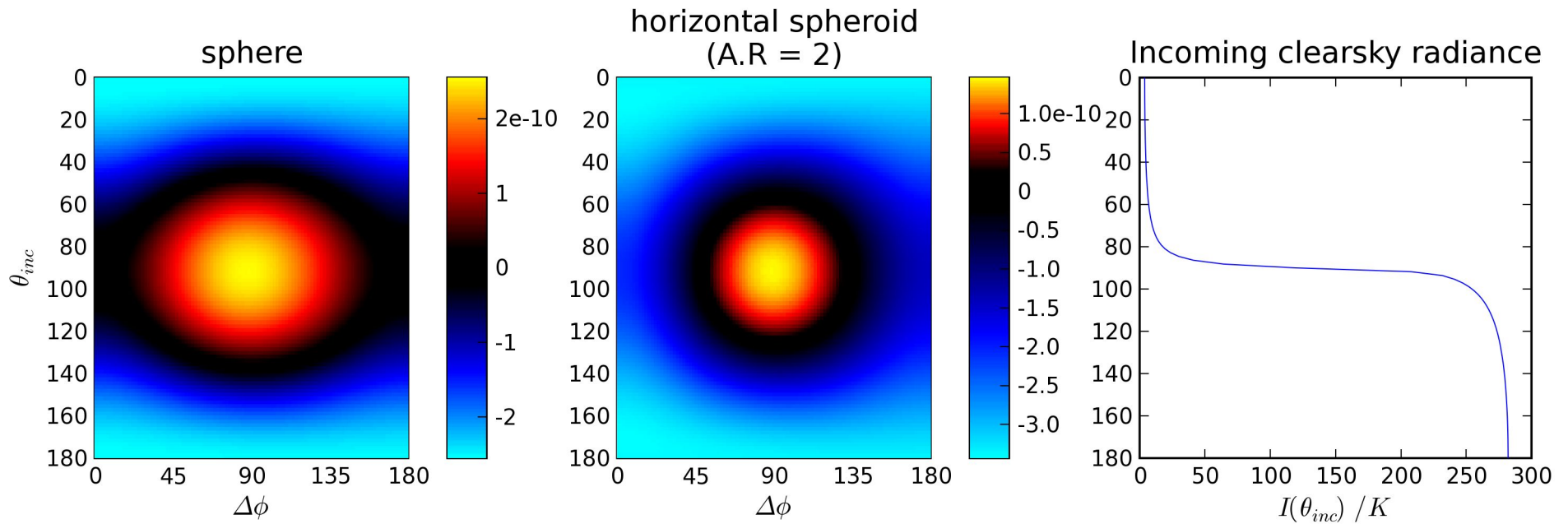
for high single scattering albedo:

$$K_{21}(\mathbf{n}) \approx \int_{4\pi} Z_{21}(\mathbf{n}, \mathbf{n}')d\mathbf{n}'$$

so the phase matrix tells all...

# $Z_{21}$

$Z_{21}(\theta_{scat}=90)$



$$K_{21} = 0$$

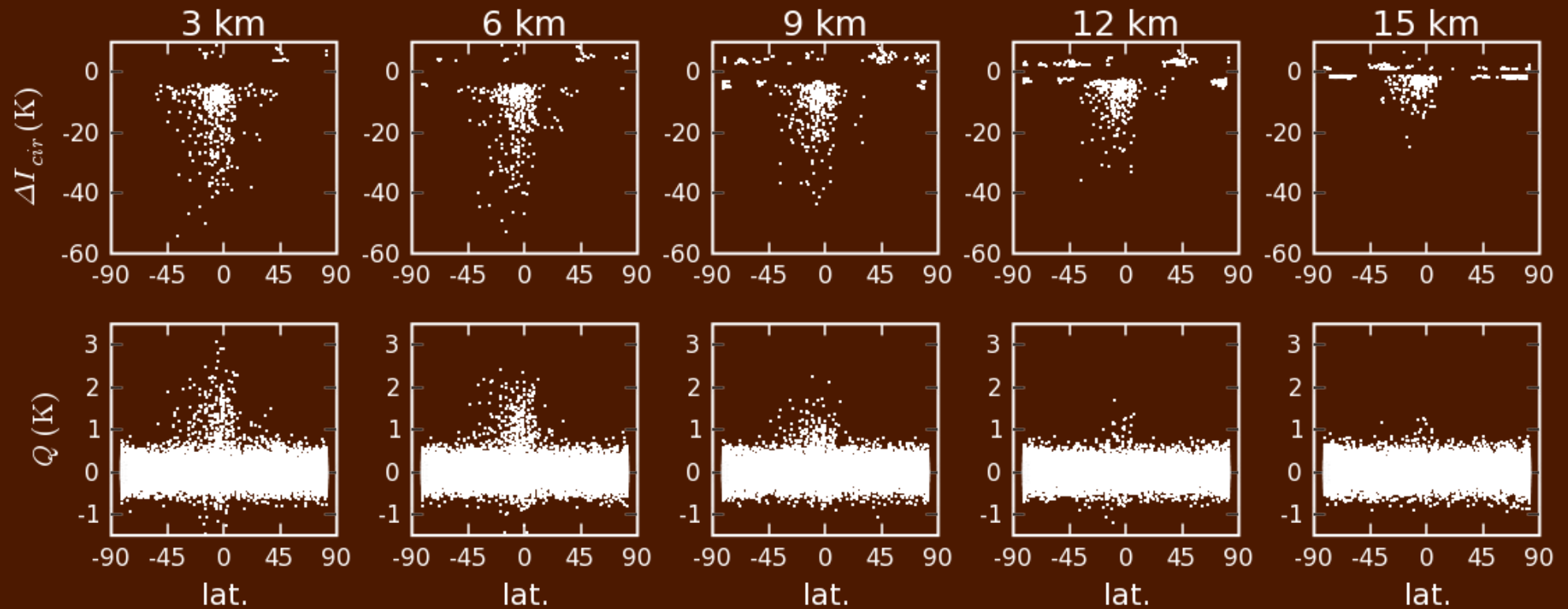
$$K_{21} < 0$$

so for low tangent heights

$$K_{21}(\mathbf{n})I_0(\mathbf{n}) \approx I_0(\mathbf{n}) \int_{4\pi} Z_{21}(\mathbf{n}, \mathbf{n}')d\mathbf{n}' > \int_{4\pi} Z_{21}(\mathbf{n}, \mathbf{n}')I(\mathbf{n}')d\mathbf{n}'$$

$$\rightarrow Q > 0$$

# What we actually observe

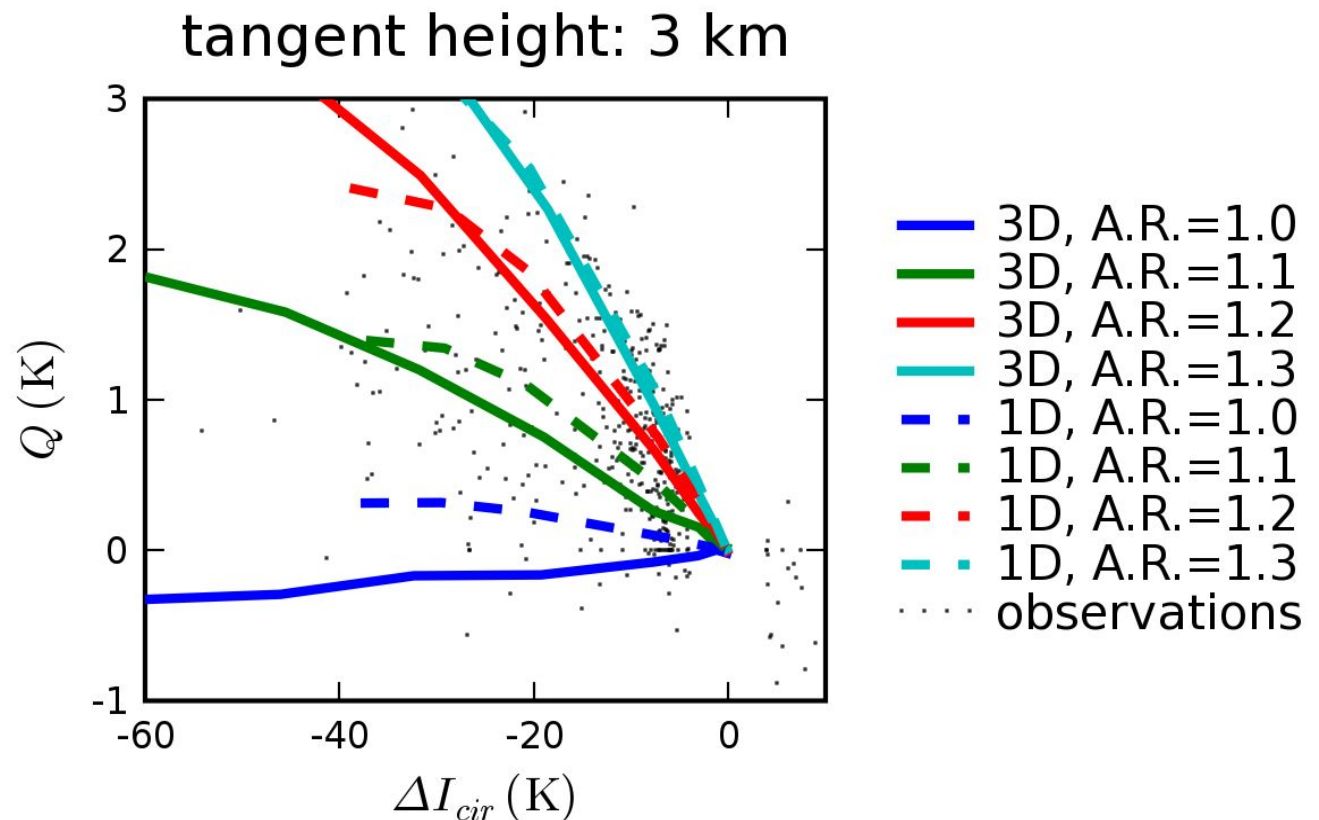




# Interpretation of Observations

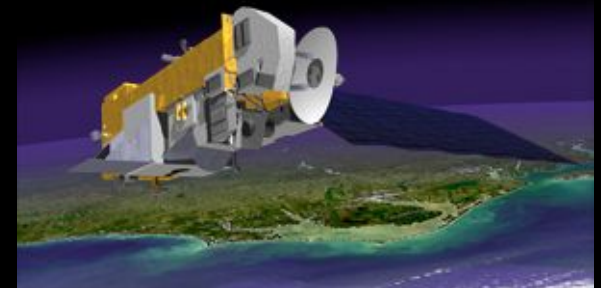
- Effect of preferential orientation is mainly determined by the ratio  $\frac{K_{12}}{K_{jj}}$
- Effect of changing the orientation distribution can be replicated by taking a single particle shape, horizontally oriented, and modifying the aspect ratio. Easiest to use oblate spheroids.

Comparison with RT simulations for 1D and 3D scenarios shows that data is consistent with a range of noticeably, but not particularly, aspherical particle shapes

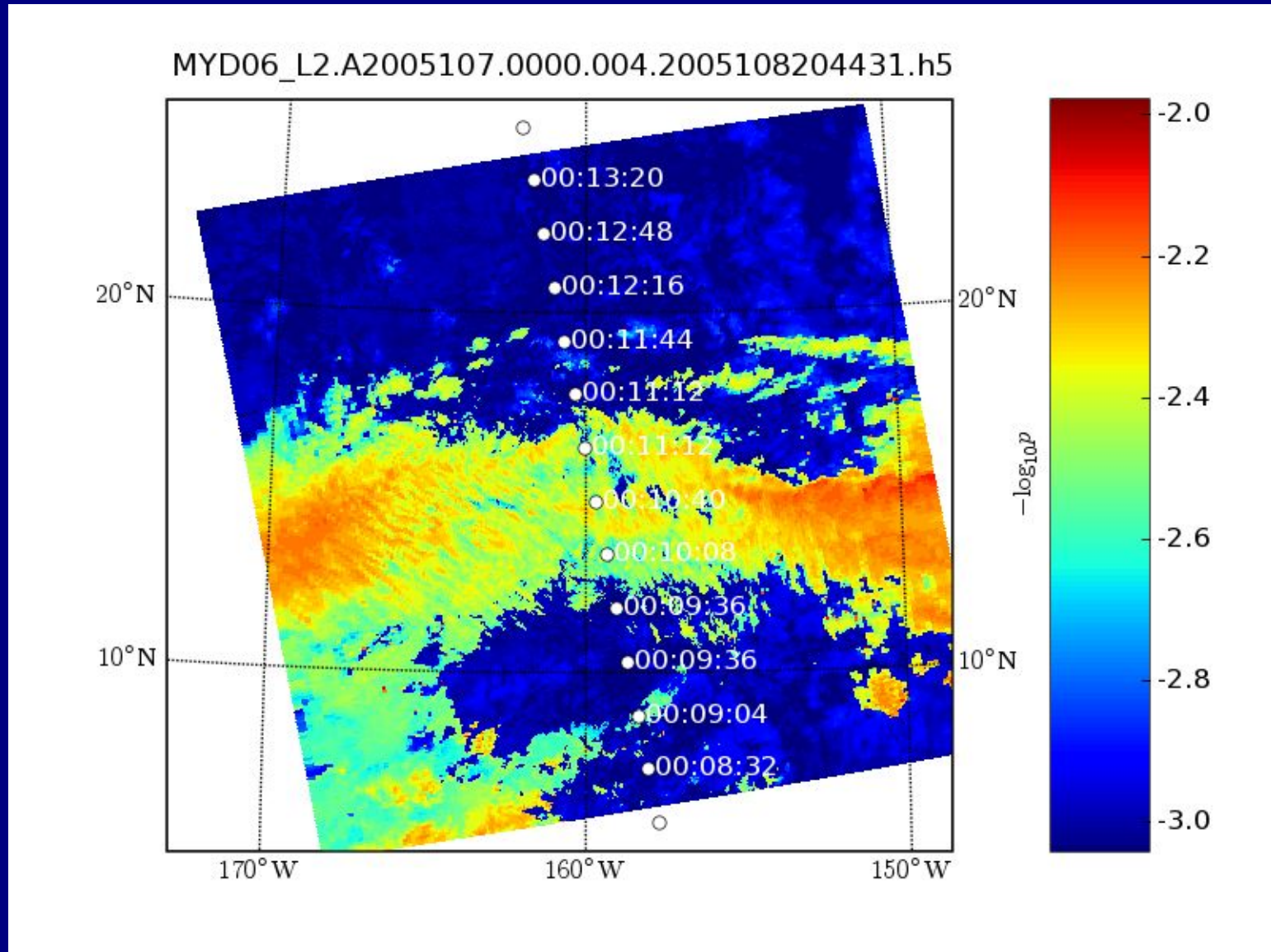


# Summary

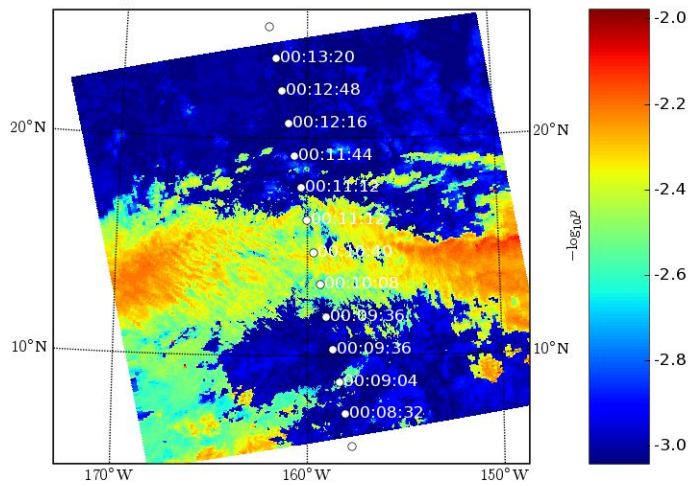
- First dual polarized microwave limb sounding observations of cirrus
- Tells us something about the degree of preferred ice crystal orientation
  - noticeable but not great – can be reproduced by using only mildly aspherical horizontally aligned particles
  - Operational assumption of random orientation seems justified
  - consequences for radiation balance?



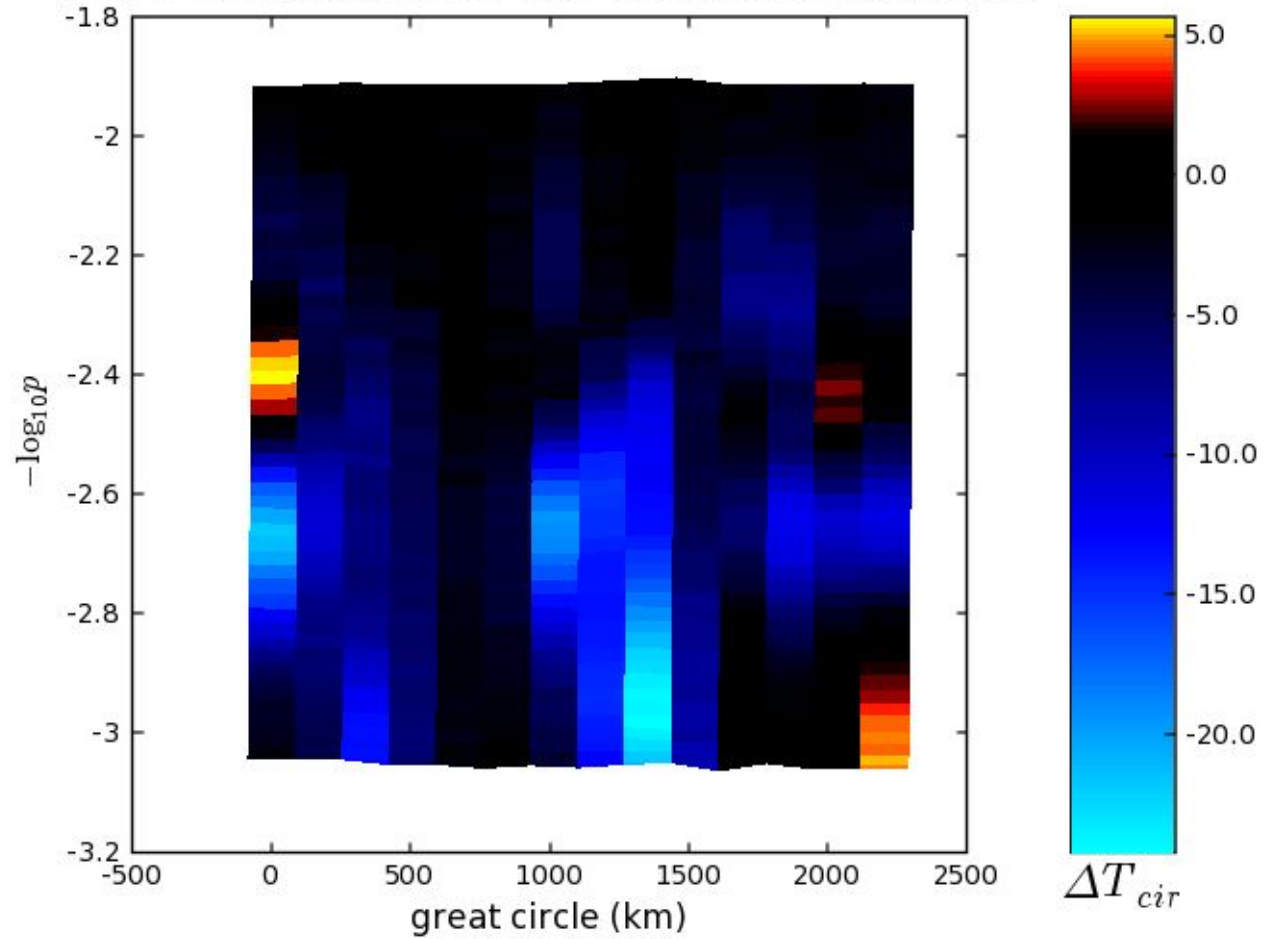
# MLS observations and MODIS cloud top heights



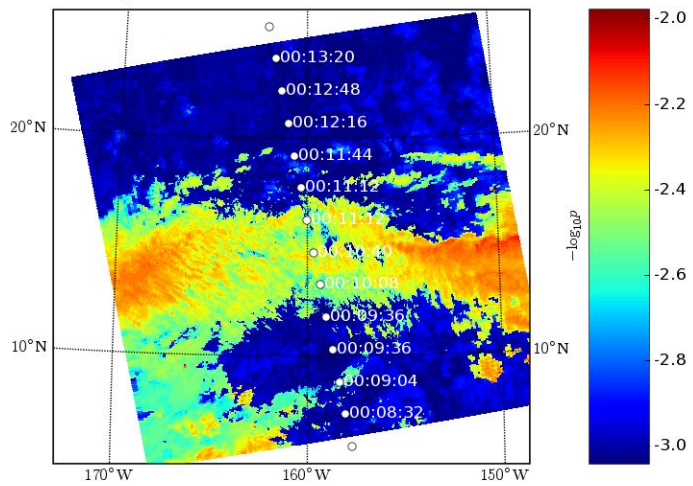
MYD06\_L2.A2005107.0000.004.2005108204431.h5



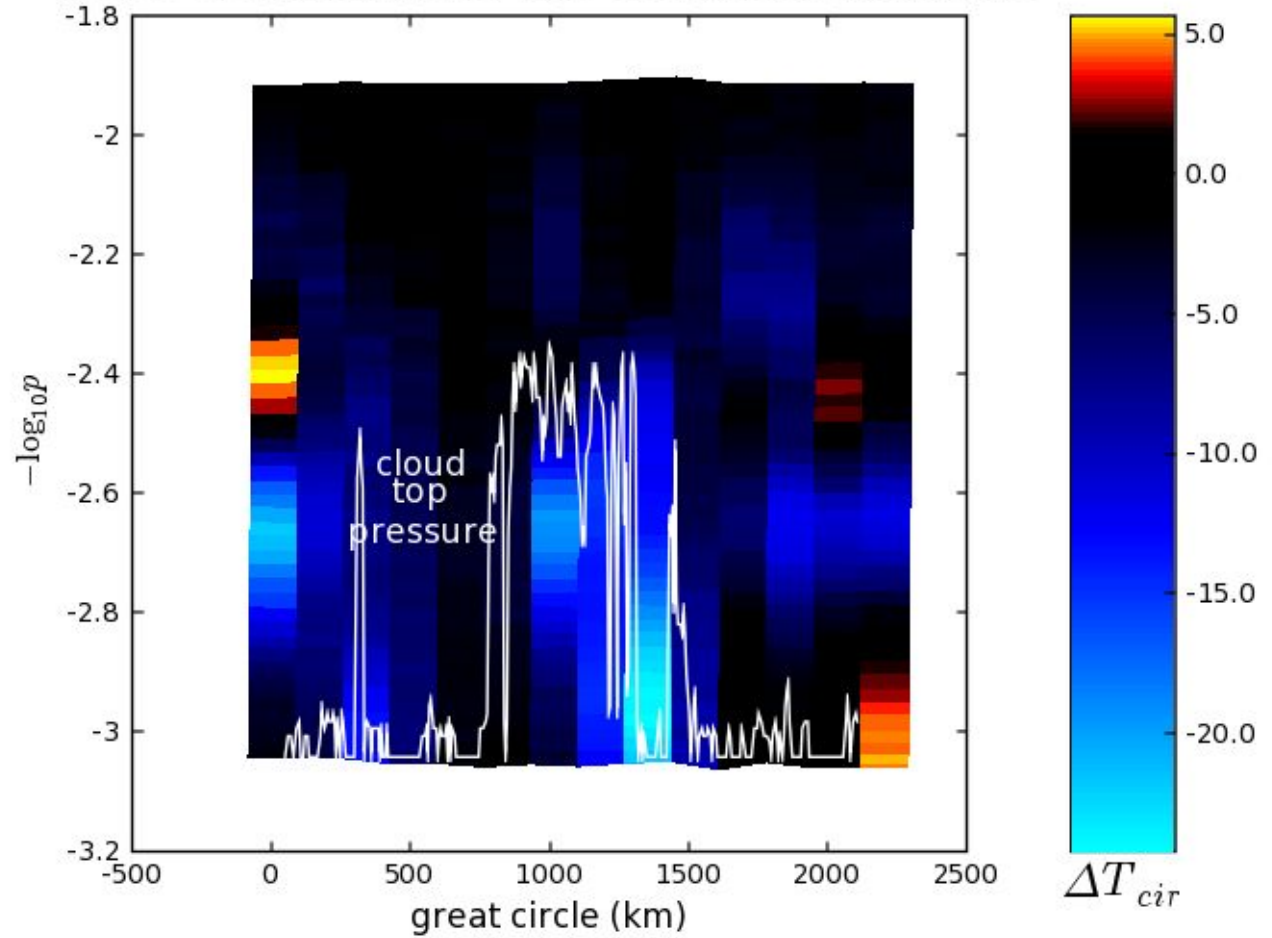
### MODIS cloud height and MLS Cloud induced radiance



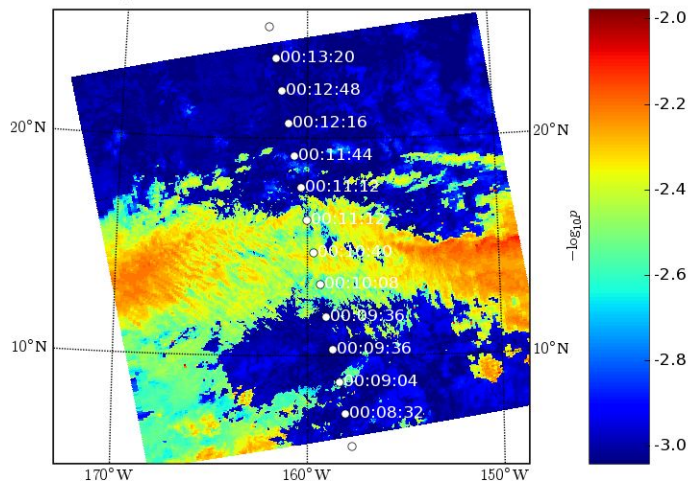
MYD06\_L2.A2005107.0000.004.2005108204431.h5



### MODIS cloud height and MLS Cloud induced radiance



MYD06\_L2.A2005107.0000.004.2005108204431.h5



### MODIS cloud height and MLS Cloud induced radiance

